

## *Observation and Modeling of Electron Cloud Instabilities*

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**Study of electron cloud effects remains very active:** 

**EPAC06: 23 abstracts mention "electron cloud"**

### ■ Over 20 publications in PRST-AB, PRL, PRA since **2003**



### *Outline*

- **Early observations, measurements**
- $\mathbb{R}^3$ Modeling development
- **EC** generation, amplification processes
- Surface science with protons, ions
- $\mathbb{R}^3$ Cures
- **Active study** 
	- $-$  Trapping in quads
	- $-$  LHC heat deposition, single bunch instability
- **New observations**
- **Summary**



### *First experimental observations*

- **Early observations of instabilities correlated with pressure attributed to** EC in **proton** rings, coasting beam or single-bunch (BINP, CERN ISR, possibly others (ZGS, AGS, Orsay, Bevatron); ~1965-72)
- Similar observations and systematic experiment study in LANL PSR **proton** ring (~1988 – today)
- **First observations in positron rings** 
	- Multibunch (KEK PF, BEPC, CESR; ~1989-97)
	- Single-bunch (KEKB, PEP-II; 1999-2000)

**[see V. Dudnikov, Proc. 2001 PAC, 1892; F. Zimmermann, PRST-AB 7, 124801 (2004)]**

- **First detailed** *in-situ* **measurements** of multipacting and EC distribution using dedicated diagnostics (RFA)
	- $-$  Positron, electron ring (APS, 1997-99)
	- $-$  Proton ring (PSR, 2000)



### *"PSR instability" story*

Evidence pointed to e-p, but many questions lead to skeptisicm

- Experimental observations first, vertical instab. (~1988)
- Data consistent with e-p theory
- The problem: Where do all the electrons come from? Why doesn't threshold change with vacuum pressure variation? How do electrons survive the gap?
- Many remained unconvinced until electron cloud measured directly with RFA, RFA sweeper (~2000)

#### RFA data lead to new understanding

- "Trailing edge multipacting" (R. Macek)
- **Proton beam loss an important source of electrons**
- **U** Very low energy electrons survive the gap without beam

Story is not finished: new questions (first turn instability)…



### *Electron cloud (EC) modeling*

Ca. 1995, **PEPII and KEKB** factories both under development; became concerned about EC effects

Calculated predictions of a **multipacting resonance in LHC**, also under development, resulted in a **crash program** at CERN to study EC effects

Modeling:

- 1<sup>st</sup> gen codes (2D analytical, PIC) developed to model EC generation and instabilities (M. Furman, K. Ohmi, F. Zimmermann, and colleagues)
- Detailed semi-empirical secondary electron emission model developed (POSINST) **[M.A. Furman, M.T.F. Pivi, PRST-AB 5, 124404 (2002)]**
- 2<sup>nd</sup> gen codes (2-3D) developed for more realistic modeling for positron, proton, heavy ion beams



# A survey of codes (incomplete)



**CMEE, Stoltz** 

Library for computational methods for electron cloud effects

**Courtesy A. Adelmann et al., ECLOUD04**

Extensive benchmarking study launched after ECLOUD02, ECLOUD04, spearheaded by F. Zimmermann Std. params for single-bunch instab: Build-up, thresh. vary by 3-100 **[E.Benedetto et al., Proc. 2004 EPAC, 2502] [see also HB2006, benchmark session]**



### *Modern study of EC effects*

Electron cloud effects have been very difficult to predict

- Surface science is complex for technical materials and accelerator environment
- **Low-energy electrons notoriously difficult to characterize** experimental uncertainties

Most advances have occurred when modeling is benchmarked against detailed measured data. Notable examples:

- APS and PSR vs. POSINST
- HCX (at LBNL) vs. WARP/POSINST
- SPS (LHC) vs. ECLOUD/HEADTAIL
- **KEKB vs. PEHT/PEHTS**
- RHIC vs. CSEC, ECLOUD, maps



### *Electron cloud generation and surface science*

- Electron cloud generation
- **I** lonization of residual gas
- Photoemission
- Secondary emission,  $\delta$ 
	- $-$  Electrons accelerated by beam
	- Beam losses, protons and ions (grazing incidence on walls, collimators)

#### Secondary EC processes

- **Electron-stimulated molecular desorption, vacuum pressure** rise/runaway (PEP-II, APS, SPS, RHIC)
- $\mathbb{R}^3$  Electron cloud trapping in magnetic fields (dipoles, quadrupoles, ion pump fringe field, etc) (HCX, PSR, CESR)
- **Interference with standard beam diagnostics (SPS)**



### *Secondary electron emission, multipacting*

- $\mathcal{L}_{\mathcal{A}}$  **Universal** δ **curve, peak values surface dependent**
- <u>ra</u> **EC lifetime depends strongly on**  $\delta_{\mathbf{0}}$  ~0.5 **(CERN, PSR)**
- <u>ra</u> **True SE distribution peaks 1-3 eV, surface independent; rediffused contrib. varies/sensitive**



**Fig. courtesy of R. Kirby**

**LHC, SPS=25ns**



Fig. courtesy of F. Ruggiero, G. Arduini

### *EC amplification processes*

Dominant source of EC can vary: KEKB vs PEPII

- **Photoemission alone can be sufficient if no antechamber (KEKB, KEK**) PF, BEPC)
- Beam-induced multipacting can lead to large amplification if  $\delta > 1$ (PEP-II, APS)

**[APS vs BEPC: K. Harkay et al., Proc. 2001 PAC, 671 (2001)]**

Multipacting condition vs. EC distribution: short bunches

- Cold-electron model [O. Gröbner, Proc. 10<sup>th</sup> HEAC, Protvino, 277, 1977]
- $\mathbb{R}^3$ Multiple kicks, energy distribution (Zimmermann, Ruggiero)
- "General" condition: dependence on EC distribution (Furman, Heifets) **[K. Harkay, R. Rosenberg, PRST-AB 6, 034402 (2003); L.F. Wang, A. Chao, H. Fukuma, Proc. ECLOUD04 (2004)]**

Trailing-edge multipacting can occur in long proton bunches



#### *General multipacting condition vs. EC distribution*

#### **APS: K. Harkay, et al., Proc. 2003 PAC, 3183; ICFA BD Newsletter 33 (2004)**



#### **L. Wang et al., ECLOUD04: RHIC. KEKB, SNS**





Figure 3: Energy distributions of the electrons at the wall and inside the beam chamber in the KEKB LER's beam. Bunch spacing is 2 ns.

#### **Modeled EC distrib; RFA agrees**

Figure 1: Electron's orbit (left column) and energy at the wall (right column). RHIC beam with bunch spacing 108 ns (top row); KEKB LER beam with bunch intensity  $3.3 \times 10^{10}$  and bunch spacing 8ns (bottom row).



Figure 2: Electron's orbit (left column) and energy at the wall (right column) in the SNS accumulator ring. Bunch length is 700 ns.

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#### *Trailing edge multipacting at Proton Storage Ring*



**Wideband coherent motion 50-300 MHz (4.4** μ**C/pulse)**





**7.7** μ**C/pulse**

**LANL Electron Sweeper RFA (~500 V pulse, 80MHz fast electronics added)**

**Prompt electron signal due to trailing-edge multipactor; swept electrons survive gap** 

**bunch length = 280 ns**

**Figs. courtesy R. Macek A. Browman, T. Wang**



#### *References and workshops*

**Review talks at Accelerator Conferences:** J.T. Rogers (PAC97), F. Ruggiero (EPAC98), K. Harkay (PAC99), F. Zimmermann, K. Harkay (PAC01), G. Arduini, F. Zimmermann (EPAC02), M. Furman, M. Blaskiewicz (PAC03), M. Pivi, L. Wang (PAC05) http://www.jacow.org

**ICFA BD Newsletter No. 33, Apr. 2004:** special edition on Electron Cloud Effects in Accelerators http://www-bd.fnal.gov/icfabd

#### **Workshops, past:**

- $\overline{\mathbb{R}}$ *Multibunch Instabilities Workshop*, KEK, 1997 KEK Proc. 97-17
- $\overline{\mathbb{R}}$  *Two-Stream ICFA Mini Workshop*, Santa Fe, 2000 http://www.aps.anl.gov/conferences/icfa/two-stream.html
- *Two-Stream Workshop*, KEK, 2001 http://conference.kek.jp/two-stream/
- $\overline{\mathbb{R}}$ *ECLOUD02*, CERN, 2002 http://slap.cern.ch/collective/ecloud02/
- $\overline{\mathbb{R}}$  *Pressure Rise Workshop*, RHIC/BNL, Dec. 2003 http://www.agsrhichome.bnl.gov/AP/PressureRise/Page1.htm
- *ICFA ECLOUD04*, Napa, CA, Apr. 2004 http://www.cern.ch/icfa-ecloud04/
- *ICFA High Brightness Hadron Beams,* KEK/JAEA, May 2006 *ftp://ftp.kek.jp/kek/abci/ICFA-HB2006*

#### *Effect of grazing incidence ions, protons*



#### 2. Non-perpendicular incident

- Beam injection and charge exchange caused beam loss are with the incident angles of mrad or less.
- At the time of AGS Booster was designed, ion desorption rate was believed to be 1 - 10
- More than 1e5 molecules can be generated per lost Au ion. The gold beam injection loss induced pressure rise has caused  $> 40\%$  loss during the acceleration at high beam intensity.
- Similarly, in early design of SNS, SEY was believed to be 0.1 - 1 per lost proton.
- SEY of proton impact is measured to be larger than 100 at grazing angles.

**Booster gold Injection** beam intensity 6.5e9 Au ions Extraction 3.2e9 Au ions anna **80 ms** Proton princing, 29MeV Au31+ princity, 1MUV/u 48 32.757  $Au31+$  $111$ Electron yield  $\frac{1}{2}$ **Serrated** Serrated surface surface าบ ਾ ਹੀ  $209 0.14 10<sup>7</sup>$  $10^7$  $\mathbf{u}$ 20 40. υū ຮັບ ū 201 **AU** EiLL BÙ. Armalı (Cirk sarı old Arnali sú trear ra b scraping scraping **SEY of proton and Au impact** P. Thieberger et al

**Electron impact gas desorption msrd at RHIC: 0.05 (0.01 after scrubbing) U. Iriso, W.Fischer, EPAC06, MOPCH134**

*S.Y.Zhang, T. Roser, Pressure-Rise Workshop Summary*





#### 4. Status and plan

- If Is not unusual that the measured desorption rates differ in orders of magnitude with similar conditions.
- Surface chemistry/physics may help for better understanding.
- It is proposed for systematic measurements according to specles, energy, charge state, and Incident angle.
- More measurements based on test stands are planned at **CERN, GSI, BNL, and others.**
- **Beam measurement In the** accelerators Is also **Important. For example, EC** Intensity threshold of 34 m long straight section in RHIC Is  $<$  60% of 17 m long chambers.



*S.Y.Zhang, T. Roser, Pressure-Rise Workshop Summary*



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#### 3. Progress in ion desorption measurement

- **Measurement at AGS** Booster, RHIC, LEAR, SPS, **LINAC3, SIS and GSI HLI** shows ion desorption rate of 10 - 1e7, under different conditions.
- The ion desorption rate of around 1e5 was measured at several accelerators.
- For low energy machine. the relevant incident angle is in mrad or less For high energy machine, it may go to urad or less
- A bunch measurement shows peak desorption rate at 87 deg.
- The adequate length of surface relevant to grazing angle measurement?

**PRST-AB 9, 063201 (2006); PRST-AB 8, 113201 (2005); PRST-AB 8, 053201 (2005); PRST-AB 7, 093201 (2004); PRA 61, 042901 (2000); PRST-AB 6, 013201 (2003)**





#### *Grooves, antigrazing surfaces (collimation)*



Figure 7. Samples with two different rectangular groove profiles, 1 mm or 5mm depth.



Figure 8. Measured SEY for a rectangular groove Cu sample at different angles. The smooth part of the sample has a  $\delta$ max=1.65.

#### **Int'l R&D Effort (SLAC, KEK, CERN, LANL, Frascati): M. Pivi et al., Proc. 2005 PAC, 24; G. Stupakov, ECLOUD04**



Figure 9. Electron dynamics in proximity of a rectangular groove surface in the presence of a dipole magnetic field. The electron is absorbed.



Figure 10. Simulated SEY for a smooth (above) and for a rectangular grooved surface (below) in a dipole field.





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#### *TiN, NEG coatings, surface roughness*



A. W. Molvik, et al., PRST-AB 7, 093202 (2004). Earlier e amis: P. Thieberger, et al., PR A 61, 042901 (2000).



P. THIEBERGER et al., PRST-AB 7, 093201 (2004)



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#### *Prototype fast beam feedback for e-p*

#### PSR/LANL, SNS/ORNL, LBNL, IU, SLAC collaboration

#### **[see R. Macek, Proc. HB2006; C. Deibele THPCH13]**

• Instability threshold rf voltage as function of feedback gain



### *Trapping in quadrupoles*

- $\overline{\mathbb{R}}$  **PSR, HCX: Large beta functions in quads, high beam loss (halo)**
- $\overline{\mathbb{R}}$  **KEKB: blowup with 3.27 bkt spacing, where are electrons (solenoids)?**
- $\overline{\mathbb{R}}$  **Self-consistent 3D modeling (WARP/POSINST), measurements [J-L Vay et al., Proc. HB2006; Proc. 2005 PAC, 525 and 1479]**
- $\overline{\mathbb{R}}$  **Cures? Clearing electrodes [L. Wang, H. Fukuma, et al., MOPLS143]**



 $-40$ 

time (µs

**AND experiment** 

#### *Proposed electron sweeper for quadrupoles (PSR)*





Snapshot of trapped electrons in a PSR quadrupole 5 μs after passage of the beam pulse. (Courtesy M. Pivi)

Schematic cross section of a proposed electron sweeping detector for a PSR quadrupole. (Courtesy R. Macek, M. Pivi)

#### **Simulation: EC at LHC dipoles** M. Furman and V. Chaplin, PRST-AB 9, 034403 (March 2006)



 $t_{\rm k} = 25$  ns

 $1.6x10^{11}$ 



- Simulation code POSINST
- -LHC arc dipole magnet
- key parameters:  $N_b$ ,  $t_b$ ,  $\delta_{\text{max}}$
- current result:  $\delta_{\text{max}}$  must be <~1.2
	- conditioning scenario needs to be formulated to achieve this
- CERN simulations:  $\delta_{\text{max}}$ ~1.3 is OK

20

15

10

5

0

 $1.0$ 

 $dF/dz$  [W/m]

• owing to simpler SEY model used



10

: cooling capacity available for EC power deposition (F. Zimmermann, LHC MAC mtg. #17 (2005))

K. HB2006 (Tsukuba, May 27-June 2, 2006)<br>R. F. H. Harman, "Electron Cloud Buildup: MI and LHC" p. 8

### *Maps*

- $\Box$ **Codes can predict 2nd order transitions**
- $\Box$  **1st order transitions, e.g. vacuum pressure rise in RHIC, cannot be modeled – physics missing**
- $\mathbb{R}^2$ **3D modeling computationally expensive**
- $\mathbb{R}^2$ **Maps proposed by U. Iriso and S. Peggs**
- $\mathbb{R}^2$  **Maps can predict 1st order transitions and identify good bunch patterns in RHIC in a fraction of computation time [for appl to LHC: T. Demma et al., THPCH047]**
- **U. Iriso, S. Peggs, PRST-AB 8, 024403 (2005);** N=N build up **Proc ECLOU04; MOPCH132, MOPCH133**  $\rho_{m\!+\!1} = q_1 \cdot \rho_m + q_2 \cdot \rho_m^2 + q_3 \cdot \rho_m^3$  $\,{}^3\!S\hskip-.7pt\gamma_n^3$ First N=0 • For a given surface, for the EC build up 3rowing N=16×10<sup>10</sup>ppb, Decaying N=00×10<sup>10</sup>ppb the only thing  $3.5$ N=0 decay (gap) $N = 10x10^{10}$  ppb  $\rho_m = \rho_{m+1}$ changing between Build-up 3  $N = 14x10''$  ppb Decay the bunch *m* and 2nd order fit  $2.5$ 3rd order fit  $First N=0$   $Fit$ **bunch** *m***+1** is ρ<sub>m</sub>  $P_{max}$  (nC/m)  $\overline{2}$ and  $\rho_\mathsf{m+1}^{}$  .  $1.5$ That is… ρ!  $\overline{1}$ • Plot  $\rho_{\mathsf{m+1}}$  vs  $\rho_\mathsf{m}$  $0.5$  $0.5$ • CSEC (shown) + ECLOUD, diff  $0.5$  $1.5$  $2.5$ 3  $3.5$  $\overline{2}$  $1.5$  $2.5$  $\rho_m$  (nC/m)  $\rho_m$  (nC/m) SEY models

### *New observations, modeling continues…*

Single bunch effects:

- KEKB: Sidebands observed due to EC, fast head-tail instab; modeling comparison good **[J. Flanagan et al., PRL 94, 054801 (2005); K. Ohmi et al., HB2006; THPCH050]**
- **LHC: Incoherent single bunch instab, blowup modeling;** potentially more a concern than heat load **[E. Benedetto et al., PRST-AB 8, 124402 (2005); Proc. 2005 PAC, 387 and 1344; THPCH018]**

Others:

- SNS: Efforts to control EC have paid off **[V. Danilov, S. Cousineau, Proc. HB2006]**
- FNAL: MI, Booster, Recycler **[R. Zwaska; V. Lebedev; A. Burov, Proc. HB2006]**
- ANL IPNS: Signature very similar to PSR, RFA data TBD **[G.E. McMichael et al., MOPCH126]**
- CESRc: ILC damping ring test bed [M. Palmer, MOPLS141]



### *Summary*

- Electron cloud effects important in high performance rings; continue to surprise us
- Much progress on cures for positron rings; recent focus on proton, ion beams
- Surface science is complex: primary, secondary effects
- Benchmarking of models against measured data is absolutely critical to advance understanding
	- $-$  RFA and variations (APS, PSR, SPS, KEKB, …)
	- GESD, gridded electron collector (HCX)
	- Other beam diagnostics: spectra, centroid, tune shift, etc
- Modeling effort driving towards massively parallel 3D
- Simplified models: maps, multipacting, impedance

■ Much work has been done: talk only touches the surface...



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G. Arduini, E. Benedetto, A. Browman, H. Fukuma, M. Furman, J. Galayda, Z. Guo, S. Heifets, U. Iriso, J.M. Jiminez, R. Kirby, A. Kulikov, R. Macek, A. Molvik, K. Ohmi, M. Pivi, R. Rosenberg, G. Rumolo, P. Thieberger, L. Wang, T-S. Wang, F. Zimmermann,…

Teams at APS, BEPC, HCX, PEPII, PSR, RHIC, SPS/LHC,…

Next workshop: ECLOUD07, early 2007, Asia organizers: K. Ohmi, H. Fukuma (KEK), E-S. Kim (PAL)

